

TWO STAR CALIBRATION ALGORITHM

This is the algorithm which lies behind the AWR initialization procedure:

putting h_1 & h_2 for the hour angle of star1 & star2

and

putting LST_1 & LST_2 for the local sidereal times at CAL1 & CAL2

and

putting RA_1 & RA_2 for the right ascensions of star1 & star2

and

putting dec_1 & dec_2 for the declinations of star1 & star2

we have:

$$h_1 = 15(LST_1 - RA_1) \quad LST_1 \text{ called from RTC}$$

$$h_2 = 15(LST_2 - RA_2) \quad LST_2 \text{ called from RTC}$$

$$h_2 - m = 15(LST_2 - RA_{2m})$$

$$c_1 = h_{2-m} - h_2 \quad c_1 \geq RA_2 - RA_{2m} \quad (LST \text{ values differ})$$

$$c_2 = dec_{2m} - dec_2$$

where m is the sidereal time interval between CAL1 & CAL2

putting:

$$w = \cos h_1$$

$$x = \sin h_1$$

$$y = \cos h_2$$

$$z = \sin h_2$$

$$s = \tan dec_1$$

$$t = \tan dec_2$$

we perform the following transformations:

$$a_1 = (tz - sx) = \tan \square_2 \sin h_2 - \tan \square_1 \sin h_1$$

$$b_1 = (sw - ty) = \tan \square_1 \cos h_1 - \tan \square_2 \cos h_2$$

$$a_2 = (y - w) = \cos h_2 - \cos h_1$$

$$b_2 = (z - x) = \sin h_2 - \sin h_1$$

from which:

$$a_1 b_2 - a_2 b_1 = (\tan \square_2 \sin h_2 - \tan \square_1 \sin h_1)(\sin h_2 - \sin h_1) - (\cos h_2 - \cos h_1)(\tan \square_1 \cos h_1 - \tan \square_2 \cos h_2)$$

& the elevation and azimuth of the polar axis relative to the true pole may be derived:

$$M_{el} = \frac{c_1 b_2 - c_2 b_1}{a_1 b_2 - a_2 b_1}$$

$$M_{az} = \frac{a_1 c_2 - a_2 c_1}{a_1 b_2 - a_2 b_1}$$